

Mössbauer effect with neutrinos and neutrino oscillations

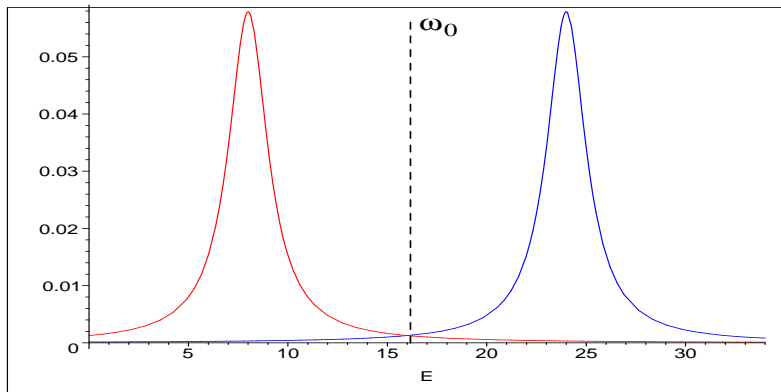
*Based on: EA, J. Kopp & M. Lindner, JHEP 0805:005,2008
[arXiv:0802.2513] and arXiv:0803.1424*

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Mössbauer effect

Conventional Mössbauer effect – Res. absorption of γ quanta:



Nuclear exc. energy: ω_0 .

Recoil energy: $R = \frac{\omega_0^2}{2M}$

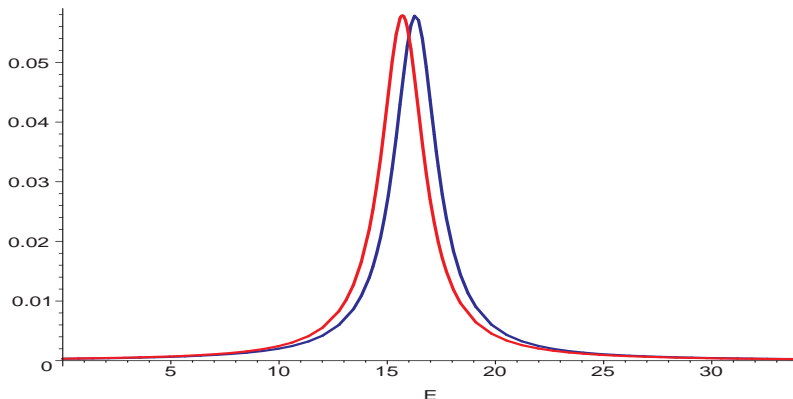
$$E_e = \omega_0 - \frac{\omega_0^2}{2M}$$

$$E_a = \omega_0 + \frac{\omega_0^2}{2M}$$

Recoilless emission and absorption (Mössb. eff.):

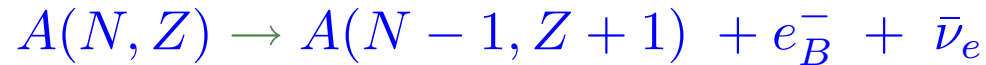
$$E_e \simeq E_a \simeq \omega_0$$

Strong enhancement of absorption

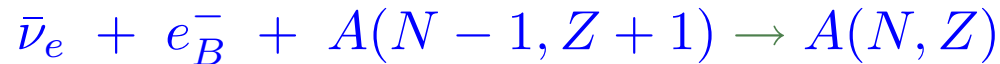


Mössbauer effect with neutrinos?

Beta decay with 2 - body final state:



Inverse process:



If the nuclei are embedded in solid state lattice, recoilless emission and absorption in principle possible.

Possibility of Mössbauer effect with neutrinos:

Visscher, 1959; Kells & Schiffer, 1983; Raghavan, 2005, 2006

Relevant processes considered:

Bahcall, 1961 – bound state β decay;

Mikaelyan, Tsinoev & Borovoi, 1967 – inverse process
(stimulated K-electron capture)

Mössbauer effect with neutrinos?

Mössbauer effect with neutrinos on ${}^3\text{H} - {}^3\text{He}$ system:



Energy release: $Q = 18.6 \text{ keV}$. Mean lifetime of ${}^3\text{H}$ is 17.8 yr \Rightarrow

Nat. linewidth $\Gamma_{{}^3\text{H}} = 1.17 \times 10^{-24} \text{ eV}$ – extremely small: $\Delta E/E \sim 10^{-28}$!

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Various (homogeneous and inhomogeneous) broadening effects exist. By suppressing them probably an effective linewidth $\Gamma_{\text{eff}} \sim 10^{-11} \text{ eV}$ can be achieved (W. Potzel) $\Rightarrow \Delta E/E \sim 10^{-15}$ – still very small.

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Very serious technical difficulties exist, but apparently realization of a Mössbauer experiment with neutrinos is not impossible (Raghavan, Potzel).

If realized: for $\Gamma \sim 10^{-11} \text{ eV}$, $\sigma \sim 10^{-33} \text{ cm}^2$!

Mössbauer effect with neutrinos?

If a Mössbauer neutrino experiment is realized \Rightarrow a unique source of extremely monochromatic low energy neutrinos. Would open up possibilities

- to detect for the first time keV neutrinos
- to detect neutrinos with g or 100 g scale (rather than t or kt scale) detectors
- to observe gravitational redshift of neutrinos
- to study neutrino oscillations at distances ~ 10 m rather than km or hundreds/thousands of km
- to search for the effects of yet unmeasured mixing angle θ_{13} and possibly measure it
- to discriminate between the normal and inverted neutrino mass hierarchies without using matter effects
- to study possible oscillations into sterile neutrino states

The purposes of our study:

Assuming that a Mössbauer experiment with ν 's can be realized

- ◇ To find out if Mössbauer neutrinos will oscillate
- ◇ To re-calculate the rate of the overall production – propagation – detection process

Will Mössbauer neutrinos oscillate?

Neutrino oscillations require some intrinsic uncertainty of energy and momentum of the emitted and detected neutrino states !

If E and p were known precisely, from $E^2 = p^2 + m_i^2$ one would determine which mass eigenstate has been emitted \Rightarrow neutrinos of different mass would not be emitted coherently.

For Mössbauer effect with neutrinos in ${}^3\text{H} - {}^3\text{He}$ system:

$$\diamond \frac{\Delta m^2}{2E} = \frac{2.5 \times 10^{-3} \text{ eV}^2}{2 \cdot 18.6 \text{ keV}} \simeq 6.7 \cdot 10^{-8} \text{ eV} \gg \Gamma \sim 10^{-11} \text{ eV}!$$

Can neutrinos of different mass be accommodated within such a small energy uncertainty?

Will neutrinos with such small energy uncertainty oscillate ?

Coherence at neutrino production

If by accurate E and p measurements one can tell which mass eigenstate is emitted, the coherence is lost and oscillations disappear!

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If $\sigma_{m^2} < \Delta m^2 = |m_i^2 - m_k^2|$ – one can tell which mass eigenstate is emitted.
 $\sigma_{m^2} < \Delta m^2$ implies $2p\sigma_p < \Delta m^2$, or $\sigma_p < \Delta m^2 / 2p \simeq l_{\text{osc}}^{-1}$.

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But: To measure p with the accuracy σ_p one needs to measure the momenta of particles at production with (at least) the same accuracy \Rightarrow uncertainty of their coordinates (and the coordinate of ν production point) will be

$$\sigma_{x, \text{prod}} \gtrsim \sigma_p^{-1} \sim l_{\text{osc}}$$

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\Rightarrow Oscillations washed out (Kayser, 1981)

How about Mössbauer neutrinos?

Very small effective linewidth $\Gamma \Rightarrow$ small energy uncertainty of the emitted neutrino state. Can different neutrino mass eigenstates be emitted coherently?

$$\sigma_{m^2} = [(2E\sigma_E)^2 + (2p\sigma_p)^2]^{1/2}$$

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– not the case. σ_X is certainly smaller than the size of the crystal (a few cm). In reality it is of the order of interatomic distances (one can destroy the crystal and find out which tritium atom turned into helium).

$$\Rightarrow \sigma_p \sim 10 \text{ keV}, \quad \text{i.e.} \quad \sigma_m^2 \simeq 2p\sigma_p \sim 4 \times 10^8 \text{ eV}^2 \gg \Delta m^2$$

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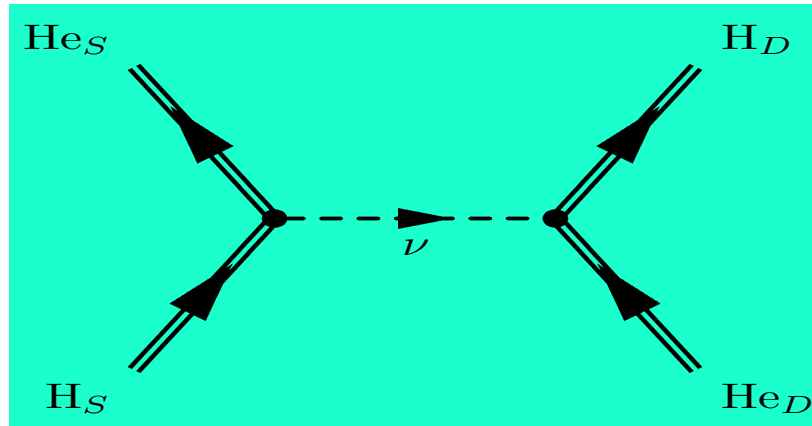
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\Rightarrow **Oscillations must occur !**

Quantum field theoretic calculations

Consider neutrino production, propagation and detection as a single inseparable process. The amplitude given by a “large” Feynman diagram:



External particles (tritium and helium) described by localized wave functions.

Neutrino is an internal line described by a propagator rather than a wave function – no assumptions on its properties (same energy, same momenta, etc.) necessary.

Model for bound tritium and helium nuclei: ground state of the corresponding harmonic oscillators (Einstein’s model of crystal) – good qualitative description of conventional Mössbauer effect. Consider $T = 0$.

QFT calculation

Inhomogeneous line broadening: Calculate the probability of the overall process for zero linewidths and then average the result over the energy distribution of ${}^3\text{H}$ and ${}^3\text{He}$ nuclei in the source and detector.

Homogeneous line broadening: modify the amplitude of the process and apply a proper averaging procedure to take into account the stochastic nature of the processes leading to homog. broadening.

⇒ Results in both cases are formally very similar. Mössbauer res. condition:

$$|E_S - E_D| \ll \gamma_S + \gamma_D$$

If it is satisfied ⇒ neutrino detection cross section enhanced by a factor

$$\sim (\alpha Z m_e)^3 / [p_e E_e (\gamma_S + \gamma_D)] \sim 10^{12}$$

compared to non-resonance $\sigma(\bar{\nu}_e + A \rightarrow A' + e^+)$ for neutrinos of same energy (assuming recoil-free fraction ~ 1).

QFT calculation – contd.

The amplitude for zero linewidths:

$$\begin{aligned}
 i\mathcal{A} = & \int d^3x_1 dt_1 \int d^3x_2 dt_2 \Psi_{He,S}^*(\vec{x}_1) e^{+iE_{He,S} t_1} \Psi_{H,S}(\vec{x}_1) e^{-iE_{H,S} t_1} \\
 & \cdot \Psi_{H,D}^*(\vec{x}_2) e^{+iE_{H,D} t_2} \Psi_{He,S}(\vec{x}_2) e^{-iE_{He,D} t_2} \\
 & \cdot \sum_j \mathcal{M}_S^\mu \mathcal{M}_D^{\nu*} |U_{ej}|^2 \int \frac{d^4p}{(2\pi)^4} e^{-ip_0(t_2-t_1)+i\vec{p}(\vec{x}_2-\vec{x}_1)} \\
 & \cdot \bar{u}_{e,S} \gamma_\mu (1 - \gamma_5) \frac{i(\not{p} + m_j)}{p_0^2 - \vec{p}^2 - m_j^2 + i\epsilon} (1 + \gamma_5) \gamma_\nu u_{e,D}
 \end{aligned}$$

Here

$$\mathcal{M}_{S,D}^\mu = \frac{G_F \cos \theta_c}{\sqrt{2}} \psi_e(R) \bar{u}_{He} (M_V \delta_0^\mu - g_A M_A \sigma_i \delta_i^\mu / \sqrt{3}) u_H \kappa_{S,D}^{1/2}$$

QFT calculation – contd.

Wave functions of bound ${}^3\text{H}$ and ${}^3\text{He}$ atoms:

$$\psi_{A,B,0}(\vec{x}, t) = \left[\frac{m_A \omega_{A,B}}{\pi} \right]^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_A \omega_{A,B} |\vec{x} - \vec{x}_B|^2 \right] \cdot e^{-iE_{A,B}t},$$

$A = \{\text{H}, \text{He}\}$, $B = \{\text{S}, \text{D}\}$ (S=source, D=detector).

$$\Gamma_0 = \frac{G_F^2 \cos^2 \theta_c}{\pi} |\psi_e(R)|^2 m_e^2 (|M_V|^2 + g_A^2 |M_A|^2) \left(\frac{E_{S,0}}{m_e} \right)^2 \kappa_S$$

$$B_0 = 4\pi G_F^2 \cos^2 \theta_c |\psi_e(R)|^2 (|M_V|^2 + g_A^2 |M_A|^2) \kappa_D$$

$$Y_S = 8(\eta_S + 1/\eta_S)^{-3}, \quad Y_D = 8(\eta_D + 1/\eta_D)^{-3}, \quad \eta_{S(D)} = \sqrt{\frac{m_H \omega_{H,S(D)}}{m_{He} \omega_{He,S(D)}}},$$

$$\sigma_p^{-2} = (m_H \omega_{H,S} + m_{He} \omega_{He,S})^{-1} + (m_H \omega_{H,S} + m_{He} \omega_{He,S})^{-1}$$

QFT calculation – contd.

For Lorentzian energy distributions of external particles:

$$\rho_{A,B}(E_{A,B}) = \frac{\gamma_{A,B}/2\pi}{(E_{A,B} - E_{A,B,0})^2 + \gamma_{A,B}^2/4}$$

($A = \{H, He\}$, $B = \{S, D\}$, $E_{A,B,0} = m_A + \frac{1}{2}\omega_{A,B}$) \Rightarrow

$$\Gamma \simeq \frac{\Gamma_0 B_0}{4\pi L^2} Y_S Y_D \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \exp\left[-\frac{(p_{jk}^{\min})^2}{\sigma_p^2}\right] \exp\left[-\frac{|\Delta m_{jk}^2|}{2\sigma_p^2}\right]$$

$$\cdot \frac{1}{2} \left(e^{-L/L_{jk,S}^{\text{coh}}} + e^{-L/L_{jk,D}^{\text{coh}}} \right) \exp\left[-i\frac{\Delta m_{jk}^2}{2\bar{E}} L\right] \frac{(\gamma_S + \gamma_D)/2\pi}{(E_{S,0} - E_{D,0})^2 + \frac{(\gamma_S + \gamma_D)^2}{4}}$$

$L_{jk,B}^{\text{coh}}$ – coherence lengths:

$$L_{jk,B}^{\text{coh}} = \frac{4\bar{E}^2}{\gamma_B |\Delta m_{jk}^2|} = \frac{\sigma_X v_g}{\Delta v_g}, \quad \sigma_X = \frac{2}{\gamma_B} \quad (B = S, D)$$

QFT calculation – contd.

Generalized Lamb – Mössbauer (Debye – Waller) factor

$$\exp \left[-\frac{p_j^2 + p_k^2}{2\sigma_p^2} \right] = \exp \left[-\frac{(p_{jk}^{\min})^2}{\sigma_p^2} \right] \exp \left[-\frac{|\Delta m_{jk}^2|}{2\sigma_p^2} \right]$$

First factor \Rightarrow suppression of emission and absorption, i.e. a generalized Lamb-Mössbauer factor, second factor \Rightarrow suppression of oscillations.

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$|\Delta m_{jk}^2| \lesssim 2\sigma_p^2 \Rightarrow$ localization condition: Spatial localization $\sigma_x \sim 1/\sigma_p$.

Oscillations would be suppressed only if $|\Delta m_{jk}^2| \gtrsim 2\sigma_p^2$.

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$|\Delta m_{jk}^2|_{\max} \simeq 2.5 \cdot 10^{-3} \text{ eV}^2$; $\sigma_p^2 \sim (10 \text{ keV})^2 \Rightarrow$

oscillations will not be suppressed.

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For realistic values of parameters – just the expected result: the rate of no-oscillation production-detection process times the standard oscillation probability (probability of $\bar{\nu}_e$ survival). Decoherence and delocalization are negligible.

Conclusions

If a Mössbauer neutrino experiment is realized – recoillessly emitted and absorbed neutrinos will oscillate

Backup slides

Factorization approach (w/o osc.)

Wave functions:

$$\psi_{A,B,0}(\vec{x}, t) = \left[\frac{m_A \omega_{A,B}}{\pi} \right]^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_A \omega_{A,B} |\vec{x} - \vec{x}_B|^2 \right] \cdot e^{-iE_{A,B}t},$$

$A = \{H, He\}$, $B = \{S, D\}$ (S=source, D=detector).

I. Separate calculation of the production rate and (no-oscillation) detection cross section

$$\Gamma_p = \Gamma_0 X_S$$

$$\Gamma_0 = \frac{G_F^2 \cos^2 \theta_c}{\pi} |\psi_e(R)|^2 m_e^2 (|M_V|^2 + g_A^2 |M_A|^2) \left(\frac{E_{S,0}}{m_e} \right)^2 \kappa_S$$

$$X_S = 8 \left(\eta_S + \frac{1}{\eta_S} \right)^{-3} e^{-\frac{p^2}{\sigma_{pS}^2}} \equiv Y_S e^{-\frac{p^2}{\sigma_{pS}^2}}$$

$$\eta_S = \sqrt{\frac{m_H \omega_{H,S}}{m_{He} \omega_{He,S}}},$$

$$\sigma_{pS}^2 = m_H \omega_{H,S} + m_{He} \omega_{He,S}.$$

Factorization approach (w/o osc.)

Restructuring of wave function of spectator electron \Rightarrow

$$\kappa_S = \left| \int \Psi_{Z=2,S}(\vec{r})^* \Psi_{Z=1,S}(\vec{r}) d^3r \right|^2$$

($= \frac{512}{729} \simeq 0.7$ in hydrogen-like approx.)

Energy spectrum $\rho(E)$ of the emitted Mössbauer neutrinos in the approximation of zero linewidth:

$$\rho(E) = \Gamma_0 X_S \delta(E - E_{S,0}).$$

Cross section of recoilless detection:

$$\sigma(E) = B_0 X_D \delta(E - E_{D,0})$$

$$X_S = Y_S e^{-\frac{p^2}{\sigma_{pS}^2}}, \quad X_D = Y_D e^{-\frac{p^2}{\sigma_{pD}^2}}$$

$$\sigma_{pS}^2 = m_H \omega_{H,S} + m_{He} \omega_{He,S} \quad \sigma_{pD}^2 = m_H \omega_{H,D} + m_{He} \omega_{He,D}.$$

Factorization approach (w/o osc.)

The overall process rate (for zero linewidths and no oscillations):

$$\Gamma = \frac{1}{4\pi L^2} \int_0^\infty \rho(E)\sigma(E) dE = \frac{\Gamma_0 B_0}{4\pi L^2} X_S X_D \delta(E_{S,0} - E_{D,0})$$

For Lorentzian energy distributions for production and detection processes:

$$\rho(E) = \Gamma_0 X_S \frac{\gamma_S/2\pi}{(E - E_{S,0})^2 + \gamma_S^2/4}, \quad \sigma(E) = B_0 X_D \frac{\gamma_D/2\pi}{(E - E_{D,0})^2 + \gamma_D^2/4}$$

Combined rate of the neutrino production, propagation and detection process:

$$\Gamma \simeq \frac{\Gamma_0 B_0}{4\pi L^2} X_S X_D \frac{(\gamma_S + \gamma_D)/2\pi}{(E_{S,0} - E_{D,0})^2 + (\gamma_S + \gamma_D)^2/4}$$

The Mössbauer resonance condition: $(E_{S,0} - E_{D,0})^2 \ll (\gamma_S + \gamma_D)^2/4$

QFT calculation – contd.

The overall process rate:

$$\Gamma = \frac{\Gamma_0 B_0}{4\pi L^2} Y_S Y_D \int_0^\infty dE_{H,S} dE_{He,S} dE_{He,D} dE_{H,D}$$
$$\cdot \delta(E_S - E_D) \rho_{H,S}(E_{H,S}) \rho_{He,D}(E_{He,D}) \rho_{He,S}(E_{He,S}) \rho_{H,D}(E_{H,D})$$
$$\cdot \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \exp \left[-\frac{2E_S^2 - m_j^2 - m_k^2}{2\sigma_p^2} \right] e^{i(\sqrt{E_S^2 - m_j^2} - \sqrt{E_S^2 - m_k^2})L}$$

σ_p – effective momentum uncertainty of the emission/absorption processes:

$$\frac{1}{\sigma_p^2} = \frac{1}{m_H \omega_{H,S} + m_{He} \omega_{He,S}} + \frac{1}{m_H \omega_{H,D} + m_{He} \omega_{He,D}},$$

An analogue of the Debye - Waller (Lamb - Mössbauer) factor:

$$\diamond \quad \exp[-(2E_S^2 - m_j^2 - m_k^2)/2\sigma_p^2] = \exp[-(p_j^2 + p_k^2)/2\sigma_p^2]$$

QFT calculation – contd.

Integrating over the coordinates and applying the Grimus-Stockinger theorem for \vec{p} - integration of the neutrino propagator:

$$\int d^3p \frac{\psi(\vec{p}) e^{i\vec{p}\vec{L}}}{A - \vec{p}^2 + i\epsilon} \xrightarrow{|\vec{L}| \rightarrow \infty} -\frac{2\pi^2}{L} \psi(\sqrt{A} \frac{\vec{L}}{L}) e^{i\sqrt{A}L} + \mathcal{O}(L^{-\frac{3}{2}})$$

to leading order in $1/L$ one gets

$$i\mathcal{A} = \frac{-i}{2L} \mathcal{N} \delta(E_S - E_D) \sum_j \exp \left[-\frac{E_S^2 - m_j^2}{2\sigma_p^2} \right] \mathcal{M}_S^\mu \mathcal{M}_D^{\nu*} |U_{ej}|^2 e^{i\sqrt{E_S^2 - m_j^2}L} \\ \cdot \bar{u}_{e,S} \gamma_\mu (1 - \gamma_5) (\not{p}_j + m_j) (1 + \gamma_5) \gamma_\nu u_{e,D},$$

where $p_j \equiv (E_S, (E_S^2 - m_j^2)^{1/2} \vec{L}/L)$. Averaging procedure:

$$\mathcal{P} = \int_0^\infty dE_{H,S} dE_{He,S} dE_{He,D} dE_{H,D} \\ \cdot \rho_{H,S}(E_{H,S}) \rho_{He,D}(E_{He,D}) \rho_{He,S}(E_{He,S}) \rho_{H,D}(E_{H,D}) \overline{|\mathcal{A}|^2}$$

Natural linewidth dominance

Time dependence of the wave functions of unstable ${}^3\text{H}$ in the source and detector have to be modified:

$$\Psi_{H,S}(\vec{x}_1)e^{-iE_{H,S}t_1} \rightarrow \Psi_{H,S}(\vec{x}_1)e^{-iE_{H,S}t_1 - \frac{1}{2}\gamma t_1}$$

$$\Psi_{H,D}(\vec{x}_2)e^{-iE_{H,D}t_2} \rightarrow \Psi_{H,D}(\vec{x}_2)e^{-iE_{H,D}t_2 - \frac{1}{2}\gamma(T-t_2)}$$

($\gamma = (17.8 \text{ yr})^{-1} = 1.17 \times 10^{-24} \text{ eV}$ – natural linewidth of tritium).

Time integrations in the interval $[0, T]$ rather than $(-\infty, \infty)$.

QFT calculation – contd.

The result:

$$\mathcal{P} = \frac{\Gamma_0 B_0}{4\pi L^2} Y_S Y_D \frac{2}{\pi} \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \exp \left[-\frac{(p_{jk}^{\min})^2}{\sigma_p^2} \right] \exp \left[-\frac{|\Delta m_{jk}^2|}{2\sigma_p^2} \right] \\ \cdot e^{-L/L_{jk}^{\text{coh}}} e^{-\gamma T} e^{-i\frac{\Delta m_{jk}^2}{2E} L} \frac{\sin^2 \left[\frac{1}{2}(E_{S,0} - E_{D,0})T \right]}{(E_{S,0} - E_{D,0})^2}$$

Coherence length:

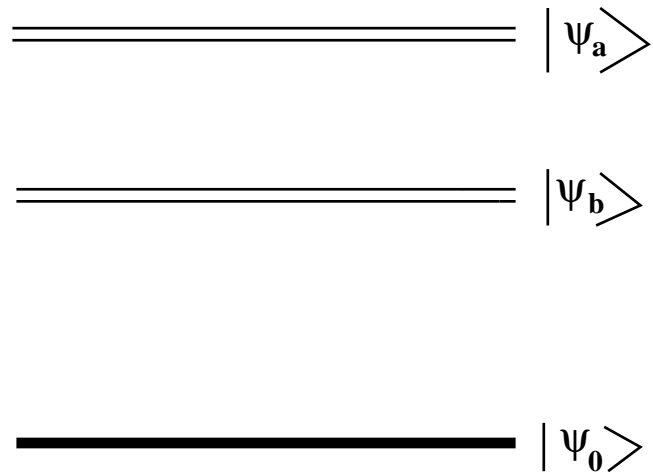
$$\frac{1}{L_{jk}^{\text{coh}}} = \gamma \left| \frac{1}{v_j} - \frac{1}{v_k} \right| \simeq \gamma \frac{|\Delta m_{jk}^2|}{2E^2}$$

Allowed detuning of $E_{S,0}$ and $E_{D,0}$ is determined not by γ but by the overall observation time T through the relation

$$|E_{S,0} - E_{D,0}| \lesssim 1/T$$

Quant. optics analogy – subnat. spectroscopy

Consider a system with two excited states, $|\psi_A\rangle$ and $|\psi_B\rangle$, with widths γ_A and γ_B .



At time $t = 0$ state $|\psi_B\rangle$ is instantaneously excited from the ground state by a short laser pulse. Also: the atom is continuously exposed to electromagnetic radiation of a frequency ω . Probability of transition into state $|\psi_A\rangle$

$$\mathcal{P} \propto \frac{1}{(E_a - E_b - \omega)^2 + (\gamma_A - \gamma_B)^2/4}$$

\Rightarrow depends on $(\gamma_A - \gamma_B)$ rather than $(\gamma_A + \gamma_B)$. In our case $\gamma_A = \gamma_B = \gamma$.

Time – energy uncertainty relation

Bilenky, v. Feilitzsch & Potzel, arXiv:0803.0527: neutrino oscillations in Mössbauer neutrino experiments contradict time-energy uncertainty relation.

Mandelstam - Tamm time-energy uncertainty relation:

$$\Delta E \Delta O \geq \frac{1}{2} \left| \frac{d}{dt} \overline{O}(t) \right|$$

$O(t)$ – an arbitrary operator in the Heisenberg representation,

$\overline{O}(t) = \langle \psi_H | O(t) | \psi_H \rangle$. Choose $O \equiv |\nu_l\rangle\langle\nu_l| \Rightarrow \overline{O} = |\langle\nu_l|\Psi(t)\rangle|^2 = P_{ll}$.

$$\Delta E \geq \frac{1}{2} \frac{\left| \frac{d}{dt} P_{ll}(t) \right|}{\sqrt{P_{ll}(t) - P_{ll}^2(t)}}$$

Bilenky et al.: set

$$x \simeq t$$

and integrate the inequality for ΔE from 0 to $t_{1\min} \equiv 2\pi E / \Delta m^2$ (“time it takes the neutrino to reach the first oscillation maximum”).

Time – energy uncertainty

⇒ Integration gives

$$\Delta E \geq \frac{1}{\pi} \frac{\Delta m^2}{2E} \quad (\text{Note : their } \Delta E \text{ is our } \sigma_E)$$

Mössbauer neutrinos: $\Delta E \sim 10^{-11}$ eV, $\frac{\Delta m^2}{2E} \sim 10^{-7}$ eV ⇒ oscillations
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satisfied for point-like neutrinos (more generally, when the size of the neutrino wave packet σ_X is small compared to the distance x they have traveled). In reality: baselines of interest $x \sim 10$ m, but $\sigma_X \sim 1/\gamma \sim 10$ km !

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⇒ Time of arrival is not well defined ⇒ integration over time is pointless.

Without time integration, using $P_{ll}(x, t) = 1 - \sin^2 2\theta \sin^2 \phi(x, t)$,

$$\Delta E \geq |E_1 - E_2|$$

– satisfied for Mössbauer neutrinos: r.h.s. can vanish because $\sigma_p \sim 10$ keV.

Raghavan's arguments

Raghavan, hep-ph/0601079:

Even without the Mössbauer effect the absorption cross section for quasi-monochromatic neutrinos strongly resonantly enhanced.

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Consider bound-state β -decay and inverse process for free ${}^3\text{H}$ and ${}^3\text{He}$ atoms. The ratio $B\beta/C\beta \simeq 0.5\%$. Recoil energy: $R \simeq 0.06 \text{ eV} \Rightarrow$ energy separation of the emission and absorption lines $E_a - E_e = 2R \simeq 0.12 \text{ eV}$. Doppler broadening at $T = 300 \text{ K}$: $\Gamma_D \simeq 0.16 \text{ eV} \Rightarrow$ overlap of the lines $\sim 40\%$.

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Detection cross section: $\sigma \sim 10^{-42} \text{ cm}^2$ – of the same order as the cross section of $\bar{\nu}_e p \rightarrow n e^+$ for reactor $\bar{\nu}_e$'s even though the energy is ~ 150 times smaller. At $L = 10 \text{ m}$: counting rate $3 \times 10^{-2} / \text{day}$ for 100 MCi ${}^3\text{H}$ source and 1 kg ${}^3\text{H}$ target – still too small.

(1Ci = 3.7×10^{10} decays/sec.)

Mössbauer effect with neutrinos?

With Mössbauer effect and for the emission and absorption linewidths $\Gamma \sim 10^{-11}$ eV: $\sigma \sim 10^{-33}$ cm² ! With recoil-free fraction $f \simeq 0.076$:

| L | ³ H | ³ He | $\bar{\nu}_e$ capt./d | R β (T=65d) |
|------|----------------|-----------------|-----------------------|-------------------|
| 5 cm | 1kCi | 100 mg | $\sim 4 \cdot 10^4$ | ~ 40 |
| 10 m | 1Mci | 1 g | $\sim 10^3$ | ~ 10 |

(1 kCi \simeq 100 mg of ³H; R β – reverse β activity; T – activation time)

W. Potzel. Phys. Scripta T127 (2006) 85:

Raghavan's estimates may be too optimistic because they do not include effects of inhomogeneous line broadening, but experiments on Mössbauer effect with neutrinos may still be possible

W. Potzel. arXiv:0810.2170:

Additional difficulties because of variations of binding energies of ³H and ³He atoms in the lattice